

Nuclear astrophysics with TAGS

Víctor Guadilla

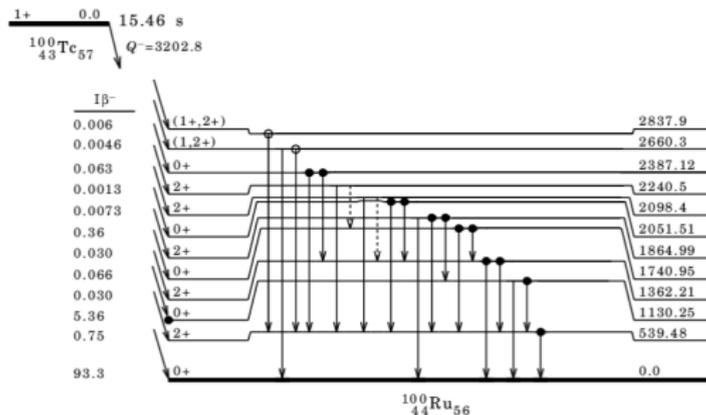
SEN group, Subatech

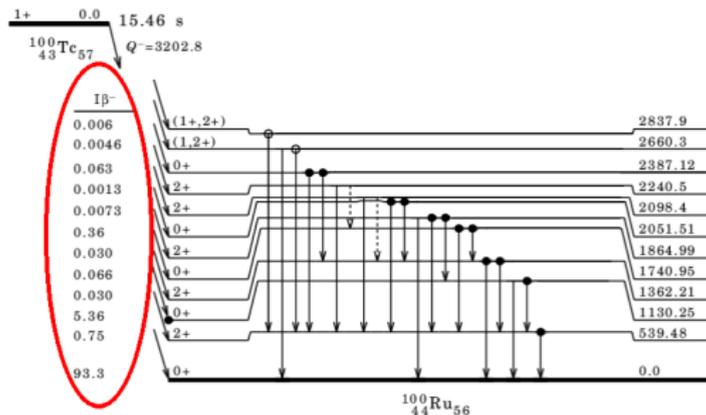


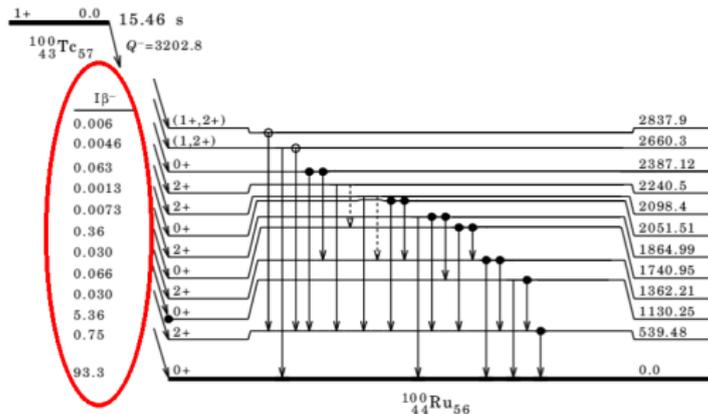
credit: Dana Berry and Erica Drezek, NASA/Goddard Space Flight Centre

- 1 Introduction: TAGS technique
- 2 Astrophysical motivation
- 3 r-process
- 4 rp-process and X-ray bursts
- 5 Summary

- 1 Introduction: TAGS technique
- 2 Astrophysical motivation
- 3 r-process
- 4 rp-process and X-ray bursts
- 5 Summary

β decay studies

β decay studies

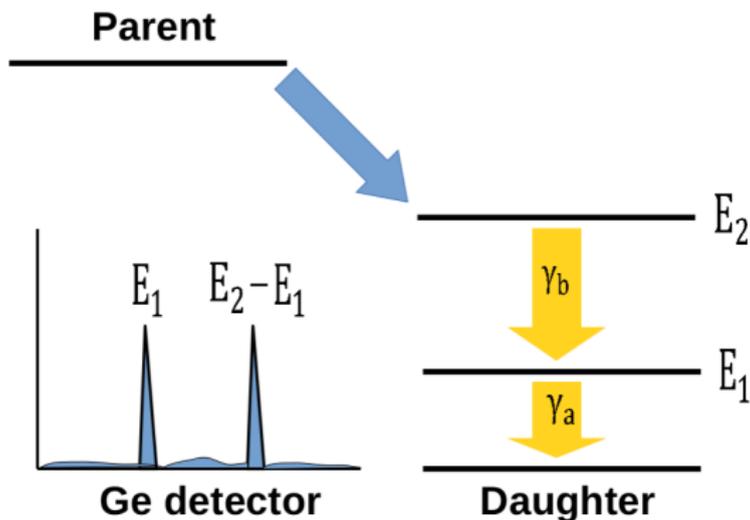
β decay studies

β -strength in the Fermi theory framework

$$\begin{aligned}
 S_{\beta}(E_x) &= \sum_{E_f \in \Delta E} \frac{\frac{1}{\Delta E} I_{\beta}(E_x)}{f(Q_{\beta} - E_x, Z) T_{1/2}} = \\
 &= \frac{1}{6146 \pm 7} \left(\frac{g_A}{g_V} \right)^2 \sum_{E_f \in \Delta E} \frac{1}{\Delta E} B(GT)_{i \rightarrow f}
 \end{aligned}$$

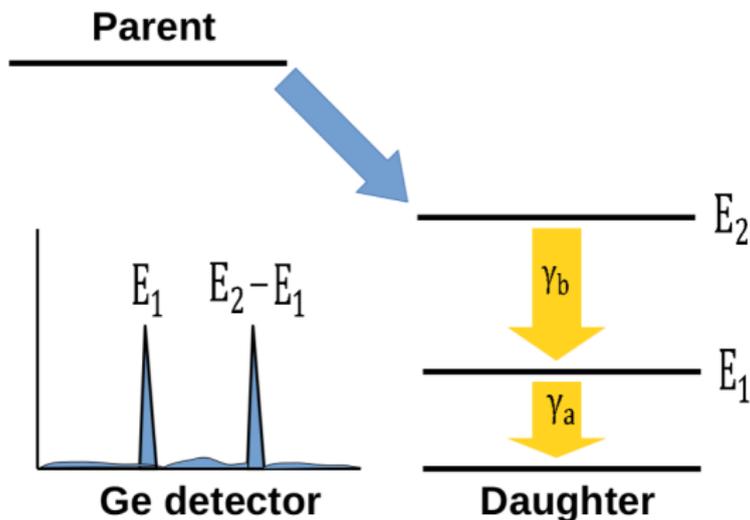
Determining I_β

I_β are often deduced from γ -intensity balance of the cascades that follow the β decay, using **HPGe detectors**:



Determining I_β

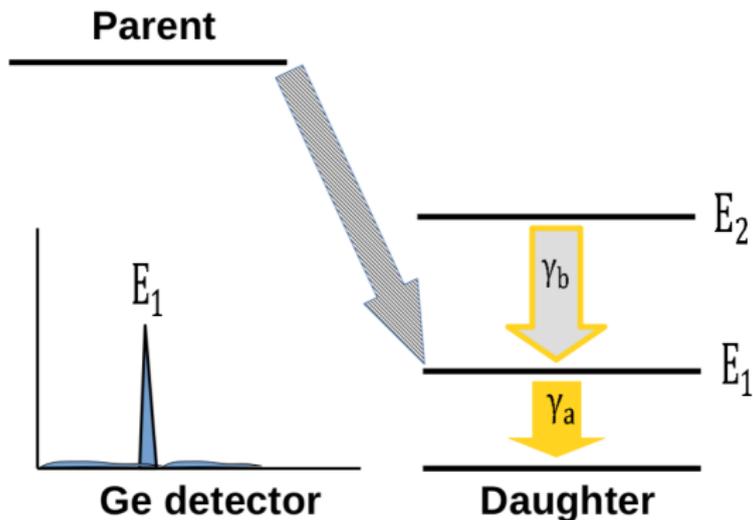
I_β are often deduced from γ -intensity balance of the cascades that follow the β decay, using **HPGe detectors**:



Low efficiency of HPGe detectors \rightarrow what happens if we miss a γ -ray?

Determining I_β

I_β are often deduced from γ -intensity balance of the cascades that follow the β decay, using **HPGe detectors**:

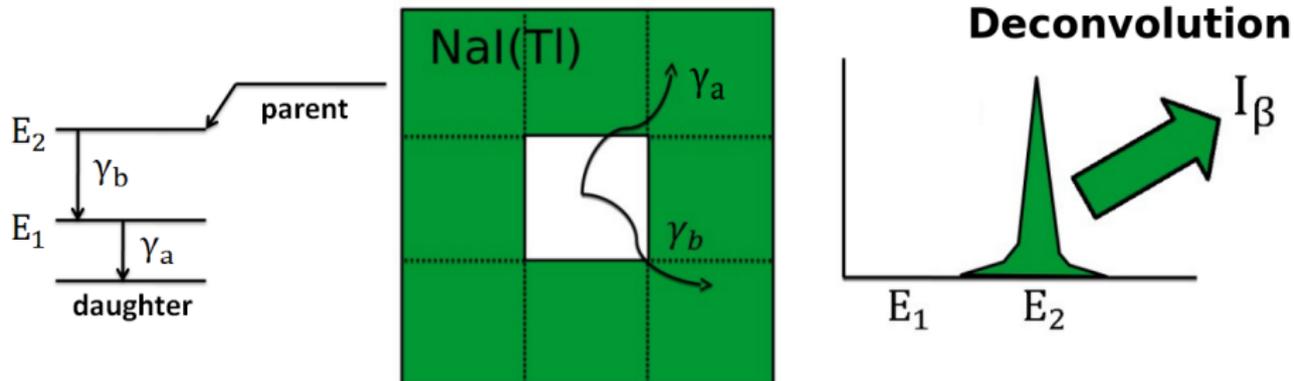


Low efficiency of HPGe detectors \rightarrow what happens if we miss a γ -ray?

Pandemonium effect J.C. Hardy et al., PLB 71 (1977) 307

Total Absorption γ -Ray Spectroscopy (TAGS)

A **T**otal **A**bsorption **S**pectrometer (TAS) acts as a **calorimeter**, absorbing the full energy released in the β -decay process.



It requires:

Large scintillator crystals covering a solid angle of $\sim 4\pi$ in order to maximize the γ -ray detection **efficiency**.

Total Absorption γ -Ray Spectroscopy (TAGS)

Inverse problem:

$$d_i = \sum_{j=1}^m R_{ij}(B) f_j$$

- $j \rightarrow$ levels, $i \rightarrow$ experimental bins
- f_j : $I_\beta(E)$ distribution
- d_i : experimental spectrum
- R_{ij} : response matrix of the detector
- B : branching ratio matrix (depends on the decay)

A deconvolution process to extract f_j

J.L. Tain and D. Cano-Ott NIMA 571 (2007) 728

Total Absorption γ -Ray Spectroscopy (TAGS)

Inverse problem:

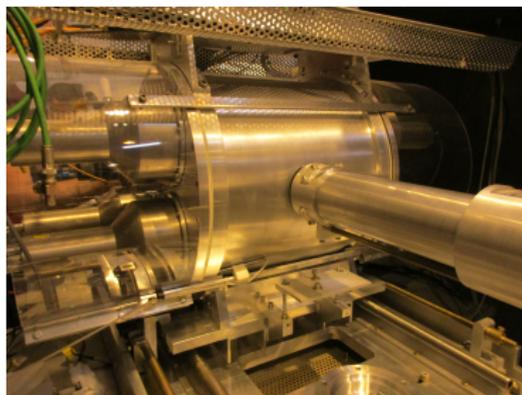
$$d_i = \sum_{j=1}^m R_{ij}(B) f_j$$

- $j \rightarrow$ levels, $i \rightarrow$ experimental bins
- f_j : $I_\beta(E)$ distribution
- d_i : experimental spectrum
- R_{ij} : response matrix of the detector
- B : branching ratio matrix (depends on the decay)

Expectation-Maximization (EM) algorithm
 A deconvolution process to extract f_j

J.L. Tain and D. Cano-Ott NIMA 571 (2007) 728

Examples of Total absorption γ -ray spectrometers I



Lucrecia



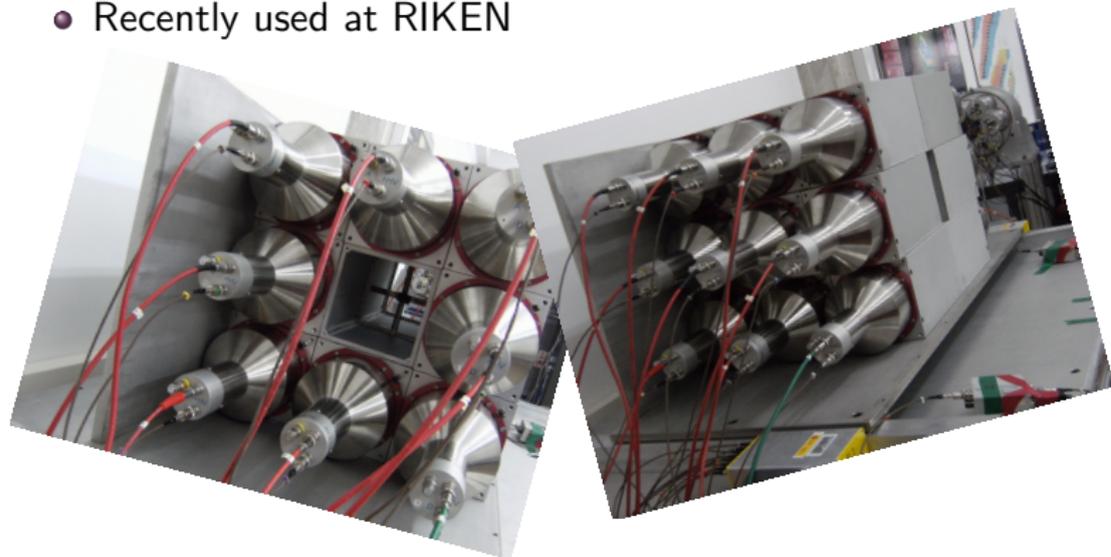
Rocinante

- **Lucrecia:** NaI(Tl) mono-crystal with cylindrical shape (38 cm diameter and 38 cm length). Permanent set-up at ISOLDE (CERN).
- **Rocinante:** cylindrical 12-fold segmented BaF₂ detector (25 cm external diameter and 25 cm length). Used in experiments at IGISOL (Jyväskylä).

Examples of Total absorption γ -ray spectrometers II

J.L. Tain et al., NIMA 803 (2015) 36

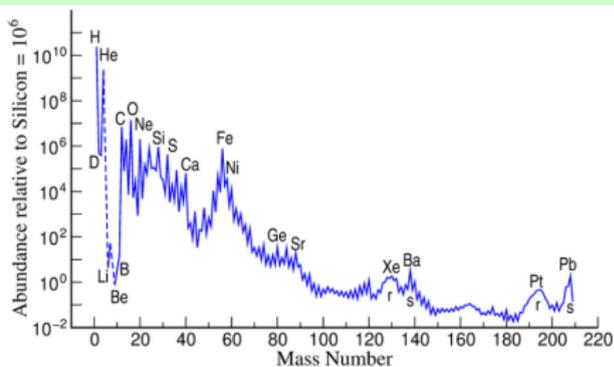
- 16-18 NaI(Tl) crystals of 150 mm \times 150 mm \times 250 mm
- Used in experiments at IGISOL (Jyväskylä)
- Recently used at RIKEN



DTAS

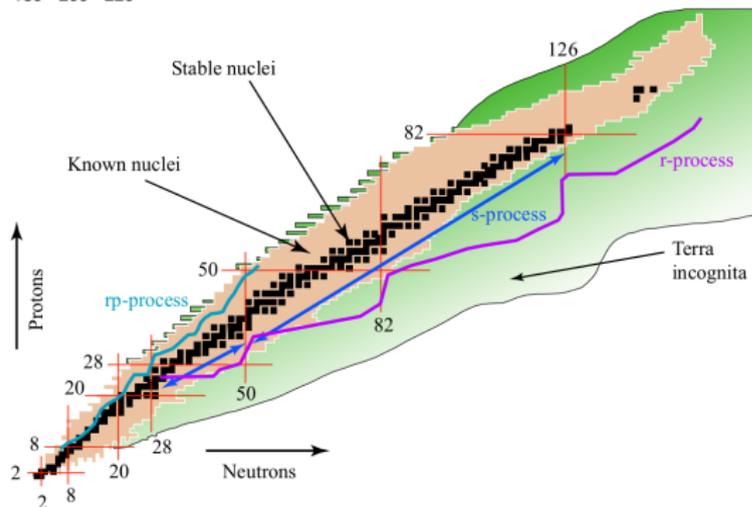
- 1 Introduction: TAGS technique
- 2 Astrophysical motivation**
- 3 r-process
- 4 rp-process and X-ray bursts
- 5 Summary

Nucleosynthesis

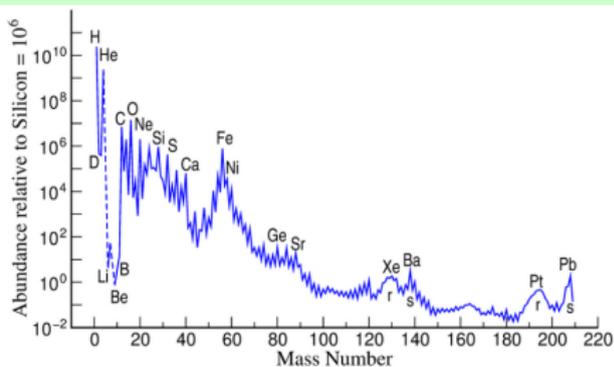


← K. Lodders Ap. J. 591 1220 (2003)

from F. X. Timmes,
Cococubed project →

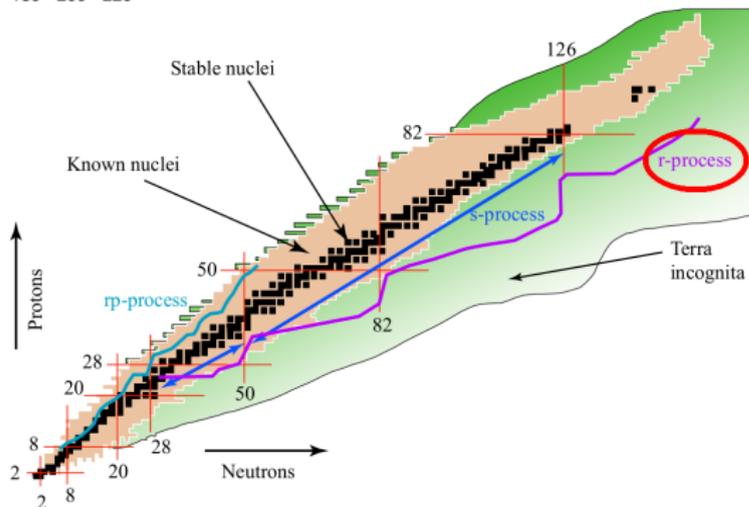


Nucleosynthesis

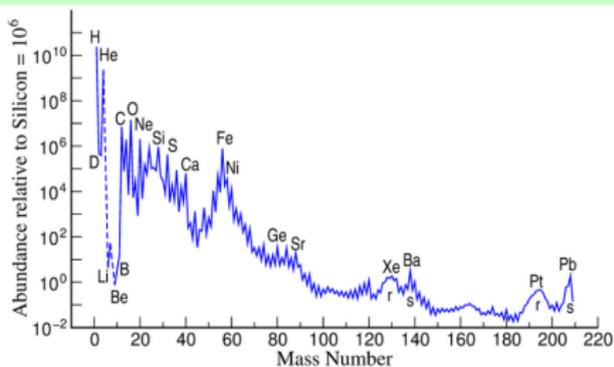


← K. Lodders Ap. J. 591 1220 (2003)

from F. X. Timmes,
Cococubed project →

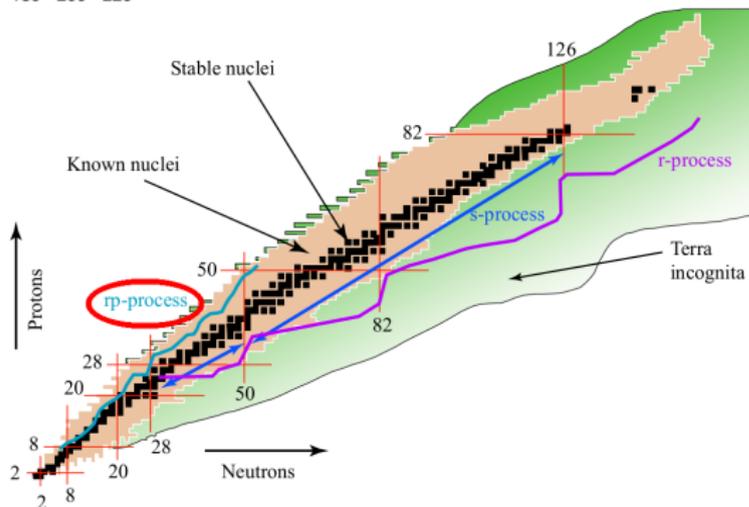


Nucleosynthesis



← K. Lodders Ap. J. 591 1220 (2003)

from F. X. Timmes,
Cococubed project →



Reproduction of the experimental abundances

Ingredients of network calculations



For exotic neutron-rich and neutron-deficient nuclei:

- Nuclear masses
- Decay rates: β^- decay, EC/ β^+ decay, p-decay, α -decay...
- Reaction rates: (n,γ) , (p,γ) , (α,γ) ...
- Fission rates and yields

Reproduction of the experimental abundances

Ingredients of network calculations



For exotic neutron-rich and neutron-deficient nuclei:

- Nuclear masses
 - Decay rates: β^- decay, EC/ β^+ decay, p-decay, α -decay...
 - Reaction rates: (n,γ) , (p,γ) , (α,γ) ...
 - Fission rates and yields
- ① From experiment, when available

Reproduction of the experimental abundances

Ingredients of network calculations



For exotic neutron-rich and neutron-deficient nuclei:

- Nuclear masses
 - Decay rates: β^- decay, EC/ β^+ decay, p-decay, α -decay...
 - Reaction rates: (n,γ) , (p,γ) , (α,γ) ...
 - Fission rates and yields
- 1 From experiment, when available
 - 2 From theory (SM, QRPA, Hauser-Feshbach...)

Reproduction of the experimental abundances

Ingredients of network calculations



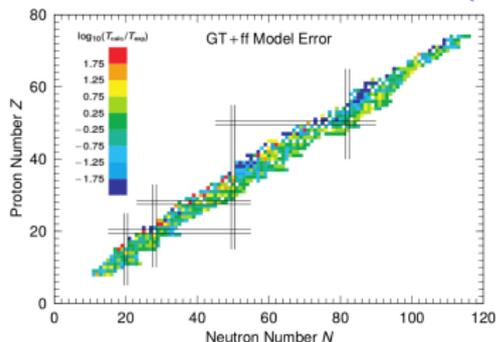
For exotic neutron-rich and neutron-deficient nuclei:

- Nuclear masses
 - Decay rates: β^- decay, EC/ β^+ decay, p-decay, α -decay...
 - Reaction rates: (n,γ) , (p,γ) , (α,γ) ...
 - Fission rates and yields
- 1 From experiment, when available
 - 2 From theory (SM, QRPA, Hauser-Feshbach...)
 - 3 From systematics

Validation of theoretical models

Validation with **integral** quantities (P_n values, $T_{1/2}$)

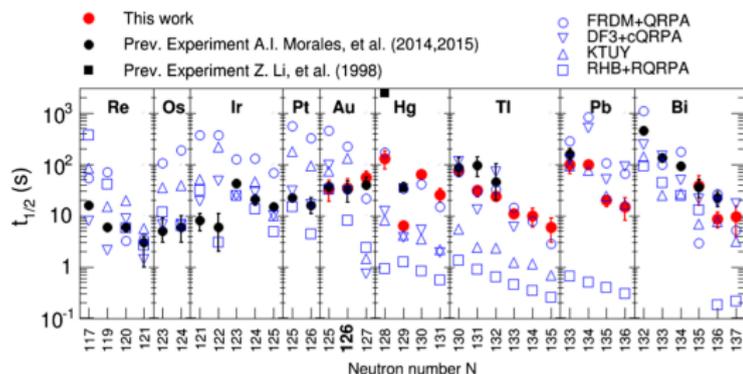
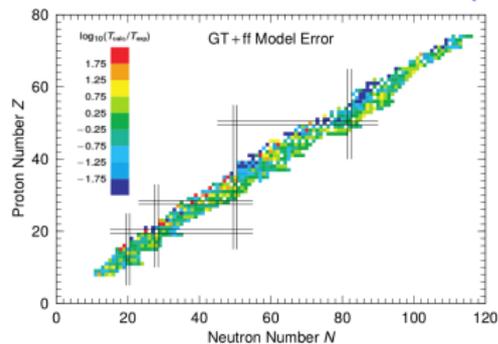
P. Möller PRC 67, 055802 (2003)



Validation of theoretical models

Validation with **integral** quantities (P_n values, $T_{1/2}$)

P. Möller PRC 67, 055802 (2003)

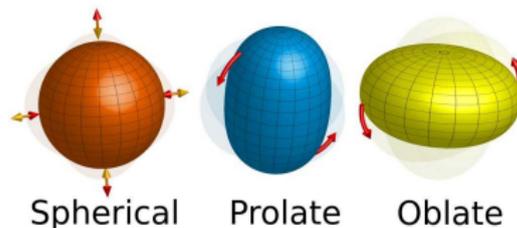
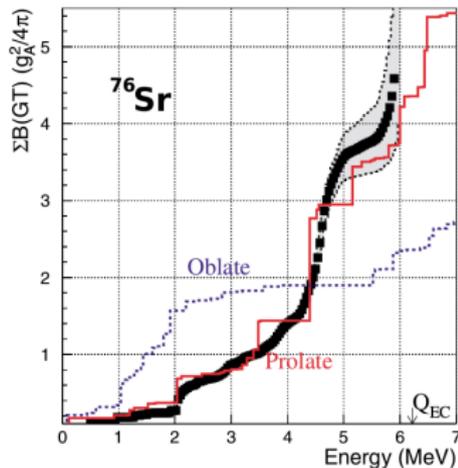


R. Caballero-Folch et al. PRL 117, 012501 (2016)

Problems to describe coherently the observed half-lives across $N=126$

Validation of theoretical models

Need of validating models with β **strength comparisons**: full information about the overlap of parent and daughter nuclear wave functions



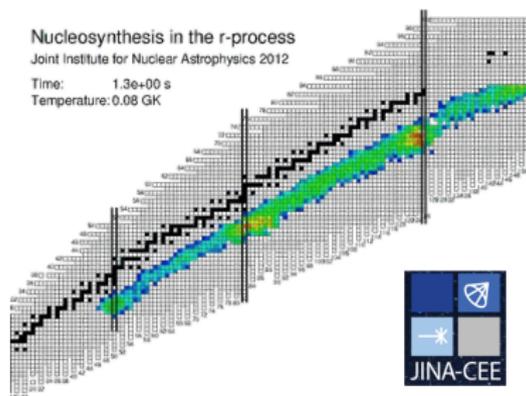
E. Nacher et al., PRL 92 (2004)
232501

QRPA calculations:

P. Sarriguren et al., PRC 89 (2014)
034311

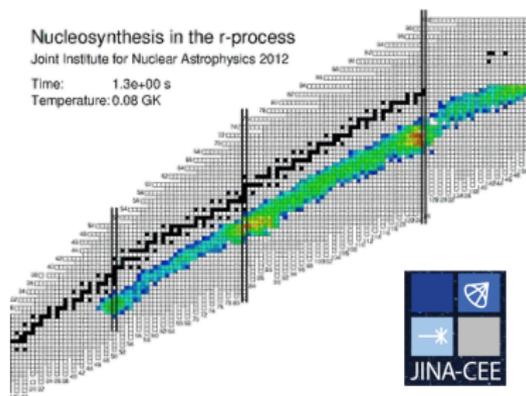
- 1 Introduction: TAGS technique
- 2 Astrophysical motivation
- 3 r-process**
- 4 rp-process and X-ray bursts
- 5 Summary

r-process



- $N_n \sim 10^{20} \text{ cm}^{-3}$ and $T \geq 1 \text{ GK}$
- Half of nuclei beyond iron
- Sites: Core Collapse Supernova, Neutron Star Mergers

r-process



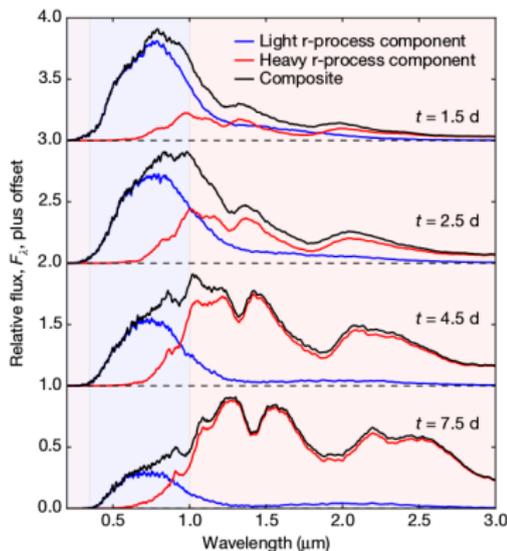
- $N_n \sim 10^{20} \text{ cm}^{-3}$ and $T \geq 1 \text{ GK}$
- Half of nuclei beyond iron
- Sites: Core Collapse Supernova, **Neutron Star Mergers**

Neutron star mergers: GW170817

Kilonova

Electromagnetic signature of the r-process due to radioactive decay of r-process nuclei

Mon. Not. R. Astron. Soc. 406, 2650–2662 (2010)



Kasen et al, Nature 551, 80 (2017)

(n,γ) reactions

- 1 Hot r-process: β decay- (n,γ) competition during freeze out
- 2 Cold r-process: equilibrium β decay- (n,γ) not reached

(n,γ) reactions

- 1 Hot r-process: β decay- (n,γ) competition during freeze out
 - 2 Cold r-process: equilibrium β decay- (n,γ) not reached
- Cross sections are not measurable: from Hauser-Feshbach (HF) statistical model calculations

(n,γ) reactions

- 1 Hot r-process: β decay- (n,γ) competition during freeze out
 - 2 Cold r-process: equilibrium β decay- (n,γ) not reached
- Cross sections are not measurable: from Hauser-Feshbach (HF) statistical model calculations
 - Parameters (NLD, PSF, NTC) for the HF calculations are obtained from data close to stability

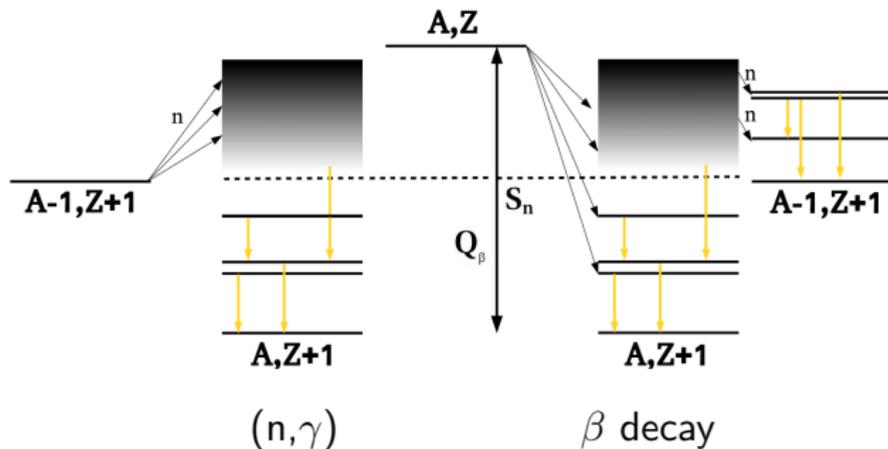
(n,γ) reactions

- 1 Hot r-process: β decay- (n,γ) competition during freeze out
 - 2 Cold r-process: equilibrium β decay- (n,γ) not reached
- Cross sections are not measurable: from Hauser-Feshbach (HF) statistical model calculations
 - Parameters (NLD, PSF, NTC) for the HF calculations are obtained from data close to stability
 - How reliable are (n,γ) HF estimations far from stability?

(n,γ) reactions

- 1 Hot r-process: β decay- (n,γ) competition during freeze out
 - 2 Cold r-process: equilibrium β decay- (n,γ) not reached
- Cross sections are not measurable: from Hauser-Feshbach (HF) statistical model calculations
 - Parameters (NLD, PSF, NTC) for the HF calculations are obtained from data close to stability
 - How reliable are (n,γ) HF estimations far from stability?
 - Constrained with different indirect techniques
[A.C.Larsen et al., Prog. in Particle and Nuclear Physics 107 \(2019\)](#)

Connection with β -delayed neutron emission



$$\sigma_{n\gamma} \propto \frac{\Gamma_\gamma \Gamma_n}{\Gamma_\gamma + \Gamma_n}$$

$$I_{\beta\gamma} \propto \frac{\Gamma_\gamma}{\Gamma_\gamma + \Gamma_n}$$

Difficulty to observe γ de-excitation from states above S_n

J.L. Tain et al., PRL 115 (2015) 062502

Some cases studied at IGISOL

Nuclide	Q_β [keV]	S_n in daughter [keV]	P_n [%]
^{87}Br	6818	5515.17	2.60
^{88}Br	8975	7053	6.58
^{94}Rb	10283	6831	10.5
^{95}Rb	9228	4345	8.7
^{137}I	6027	4025.56	7.14

2009 \longleftrightarrow Rocinante

J.L. Tain et al., PRL 115 (2015) 062502

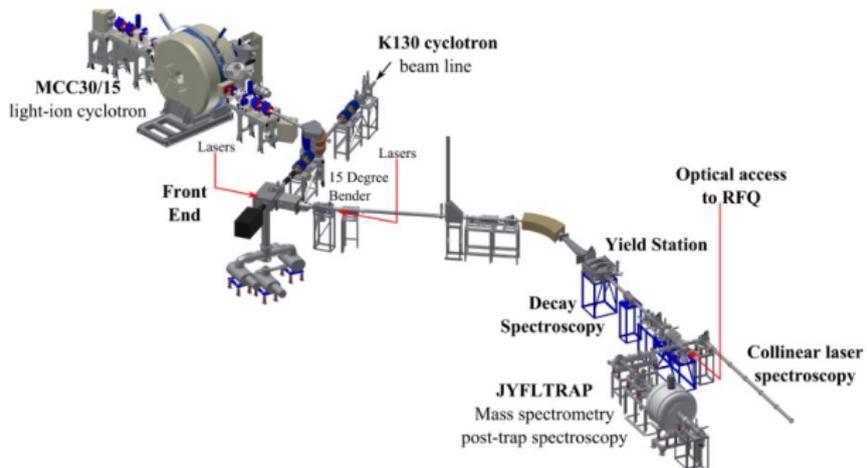
E. Valencia et al., PRC 95 (2017) 024320

2014 \longleftrightarrow DTAS

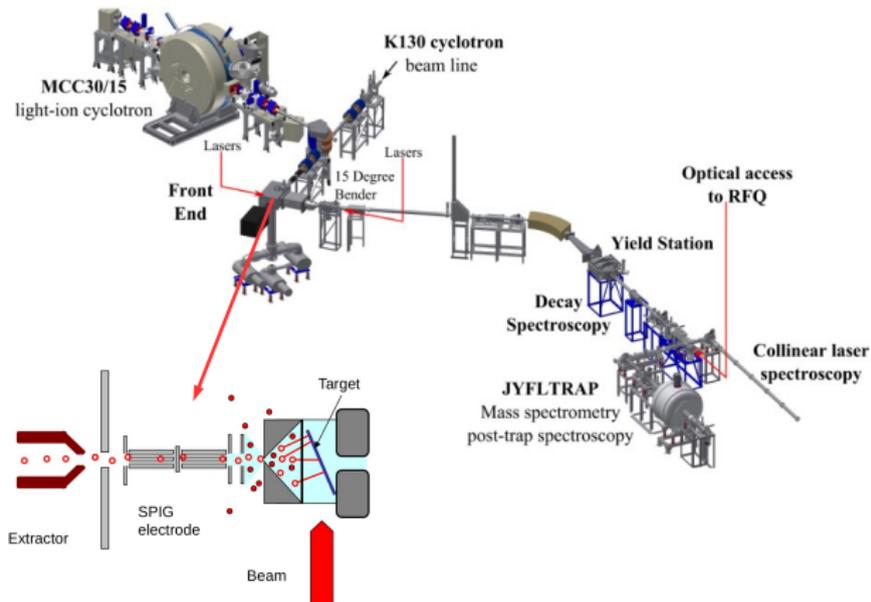
V. Guadilla et al., NIMB 376 (2016) 334

V. Guadilla et al. (2019)

IGISOL-IV: Jyväskylä (Finland)



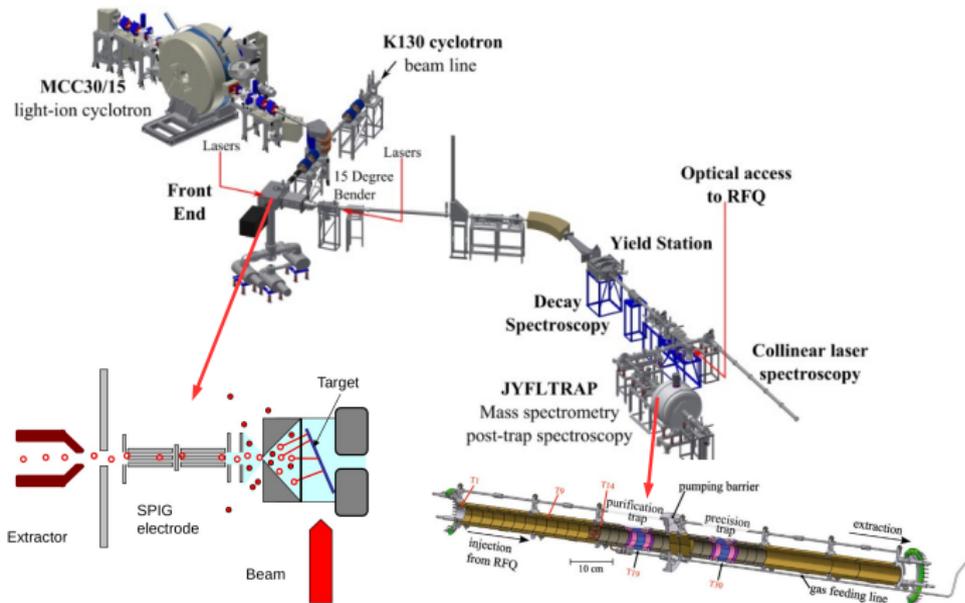
IGISOL-IV: Jyväskylä (Finland)



Fission Ion Guide

- Natural uranium target
- Refractory elements

IGISOL-IV: Jyväskylä (Finland)



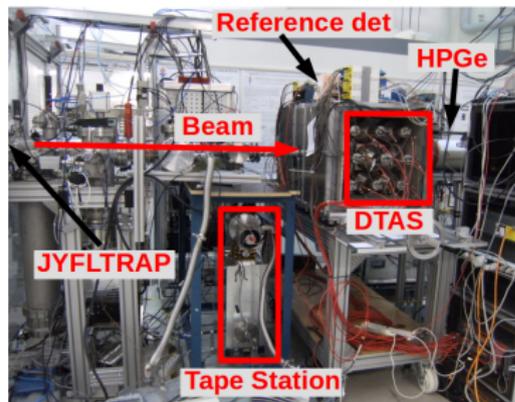
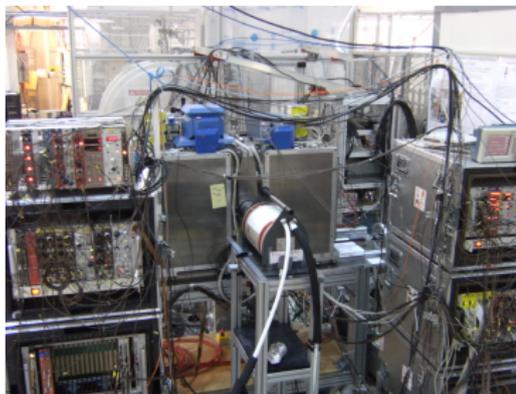
Fission Ion Guide

- Natural uranium target
- Refractory elements

JYFLTRAP Penning traps

- Inside a 7 T solenoid
- $M/\Delta M \sim 10^{5-6}$

DTAS@IGISOL: set-up



- Scintillator plastic β detector
- HPGe detector
- Tape station
- MC characterization of the detectors

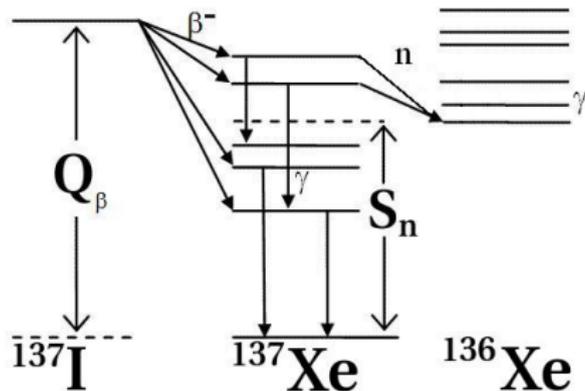


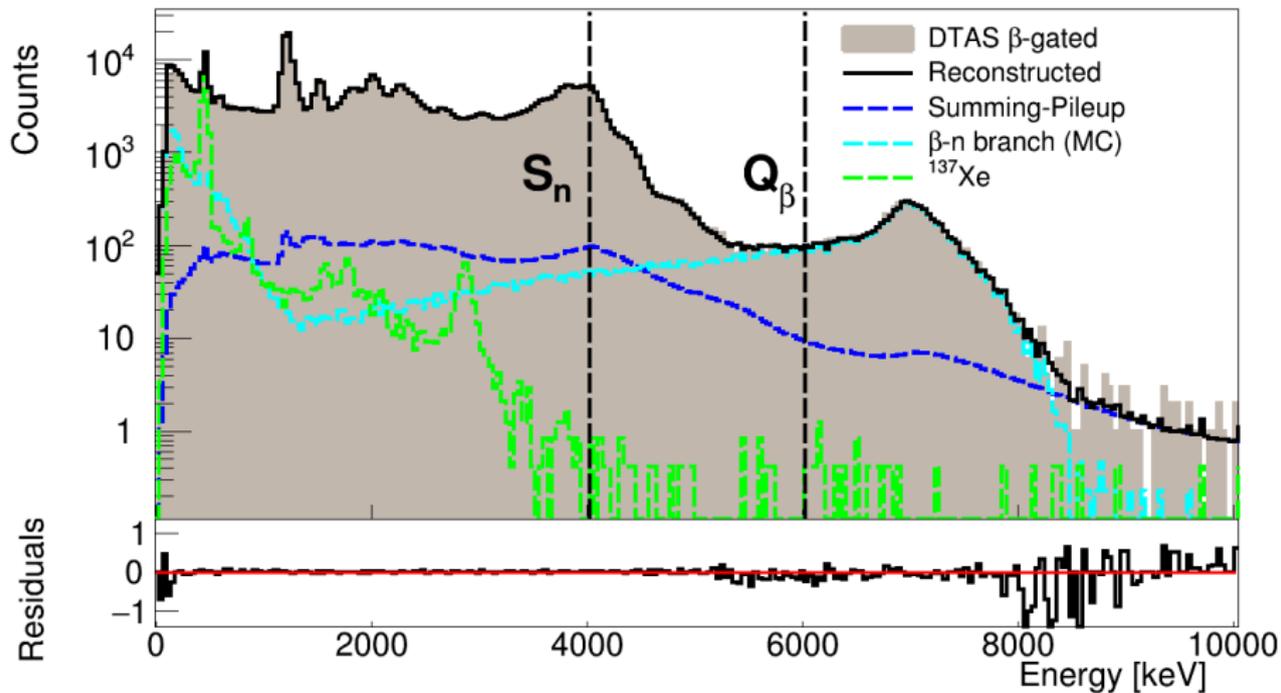
V. Guadilla et al., NIMB 376 (2016) 334

V. Guadilla et al., NIMA 910 (2018) 79

β -delayed neutron emission: ^{137}I

- $Q_\beta=6027$ keV and $S_n=4025$ keV
- $P_n=7.14\%$ and $T_{1/2}=24.5$ s
- Neutrons interact with DTAS (inelastic, capture...) \rightarrow MC



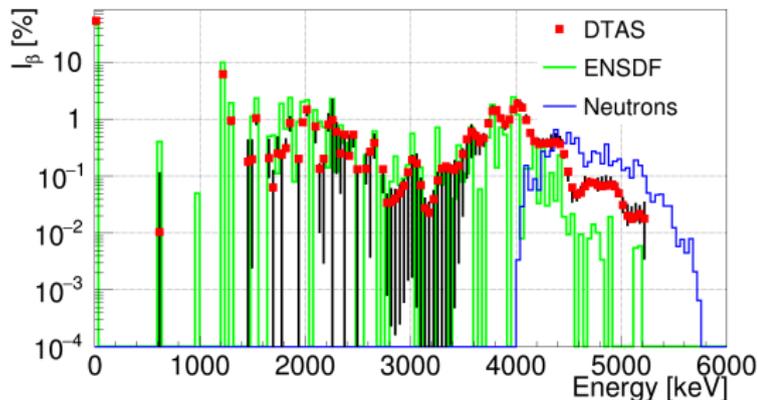
^{137}I : analysis

$$\text{Reconstructed} \rightarrow \sum_{j=1}^m R_{ij}(B) f_j(\text{final})$$

¹³⁷I: analysis

Envelope of solutions compatible with data:

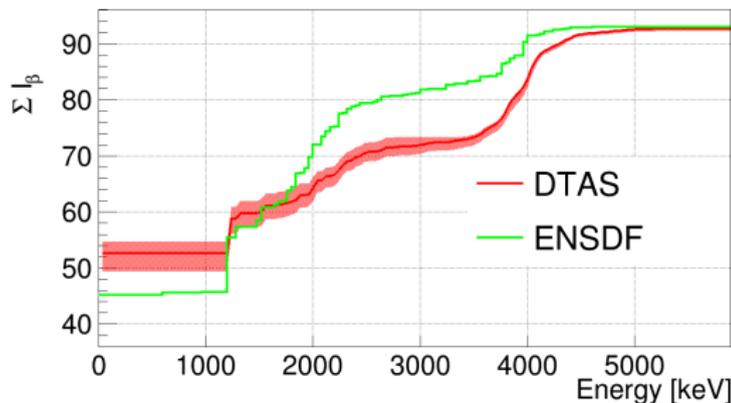
- Daughter normalization
- β -n normalization
- Summing-pileup normalization
- Known level scheme
- Deconvolution algorithm
- β detector efficiency
- β -gated **vs.** singles **vs.** neutron veto
- Branching ratio variations: parameters of level density, γ -strength function, spin-parity values of known level scheme, better reproduction of known γ -intensities etc.



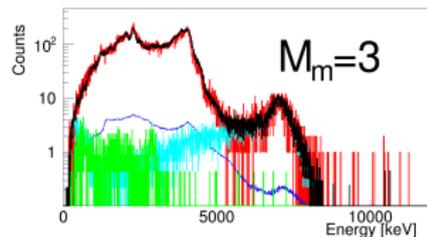
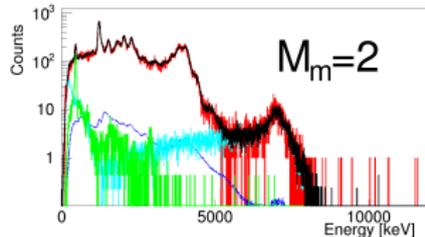
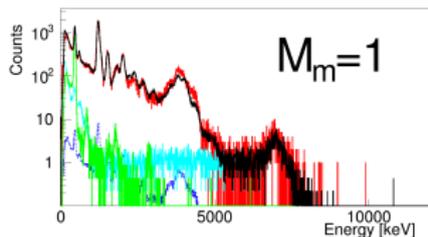
¹³⁷I: analysis

Envelope of solutions compatible with data:

- Daughter normalization
- β -n normalization
- Summing-pileup normalization
- Known level scheme
- Deconvolution algorithm
- β detector efficiency
- β -gated **vs.** singles **vs.** neutron veto
- Branching ratio variations: parameters of level density, γ -strength function, spin-parity values of known level scheme, better reproduction of known γ -intensities etc.

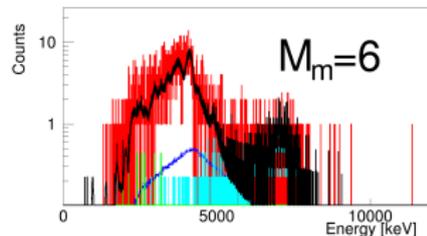
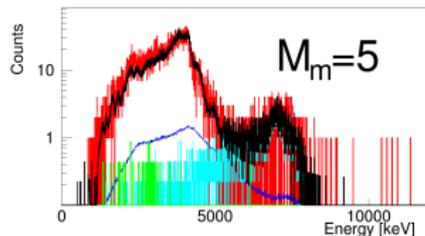
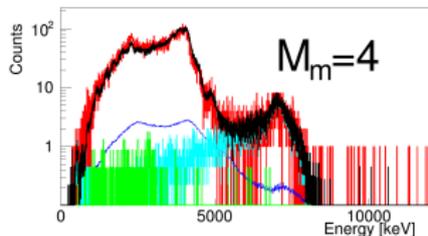


^{137}I : multiplicities

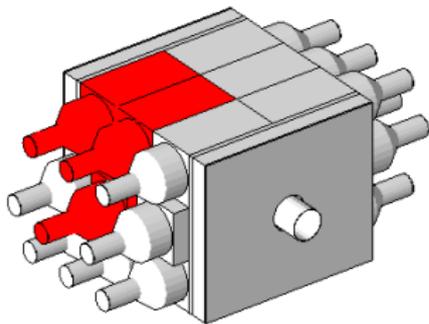


- β gated
- MC analysis
- - - Summing-Pileup
- MC neutrons
- ^{137}Xe

MC with event generator
 \leftrightarrow Input: \mathbf{I}_β of the analysis +
branching ratio matrix



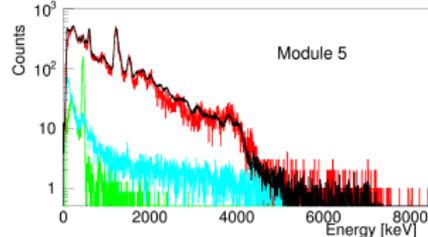
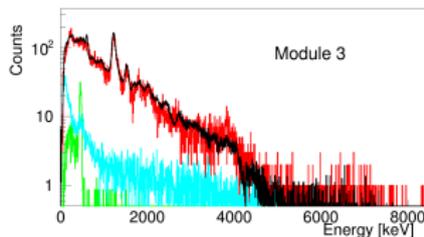
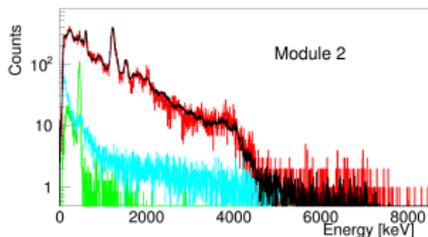
^{137}I : individual modules



Three classes of modules
geometrically equivalent

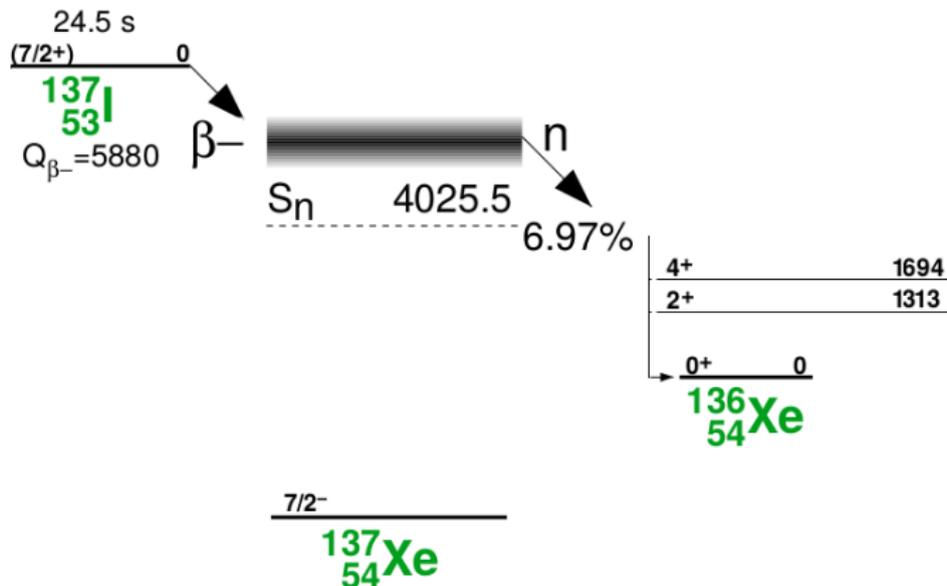
MC with event generator
 \hookrightarrow Input: \mathbf{I}_β of the analysis
 + **branching ratio matrix**

— β gated
 — MC analysis
 — MC neutrons
 — ^{137}Xe



γ emission above S_n

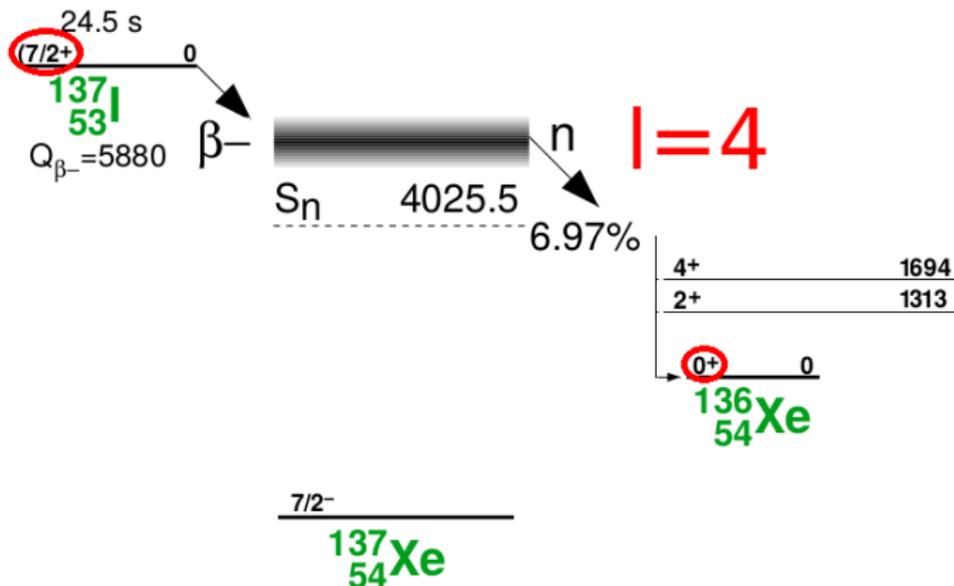
Nucleus	P_γ ENSDF [%]	P_γ DTAS [%]	P_n [%]
^{137}I	2.76	$8.88^{+1.96}_{-1.53}$	7.33(38)



V. Guadilla et al.(2019)

γ emission above S_n

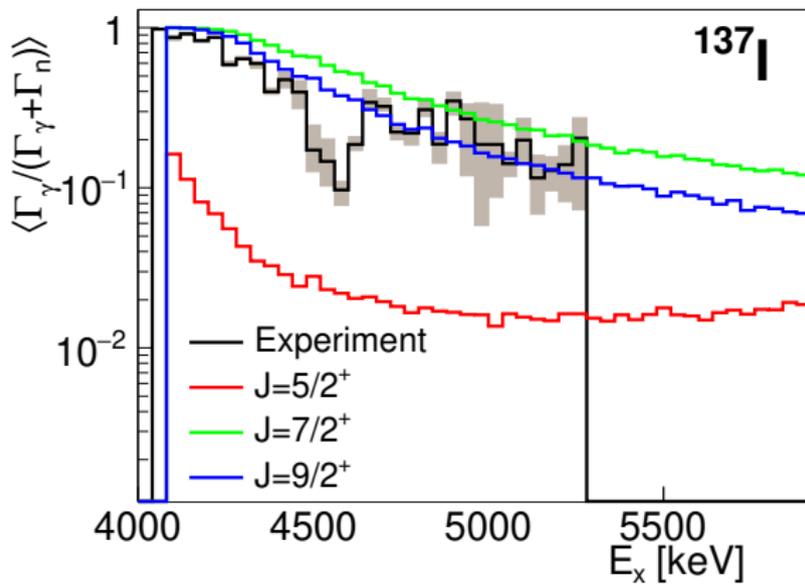
Nucleus	P_γ ENSDF [%]	P_γ DTAS [%]	P_n [%]
^{137}I	2.76	$8.88^{+1.96}_{-1.53}$	7.33(38)



V. Guadilla et al.(2019)

γ emission above S_n

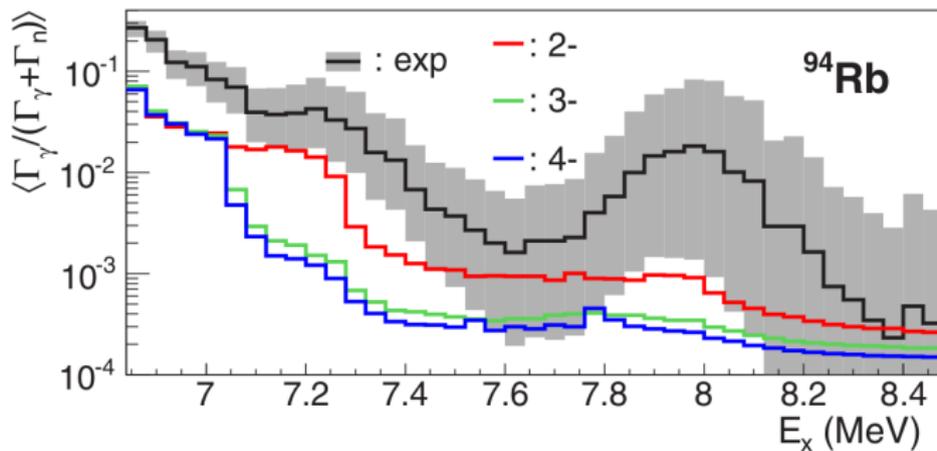
Nucleus	P_γ ENSDF [%]	P_γ DTAS [%]	P_n [%]
^{137}I	2.76	$8.88^{+1.96}_{-1.53}$	7.33(38)



V. Guadilla et al.(2019)

γ emission above S_n

Enhanced γ -branching in ^{94}Rb with respect to H-F \Rightarrow increase in the photon strength function \Rightarrow similar increase in the (n,γ) cross section

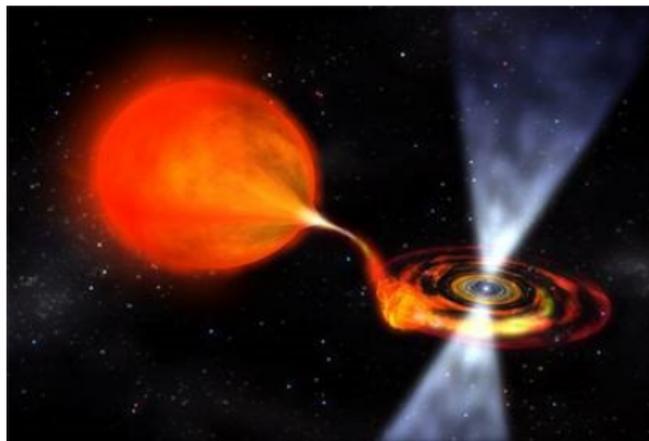
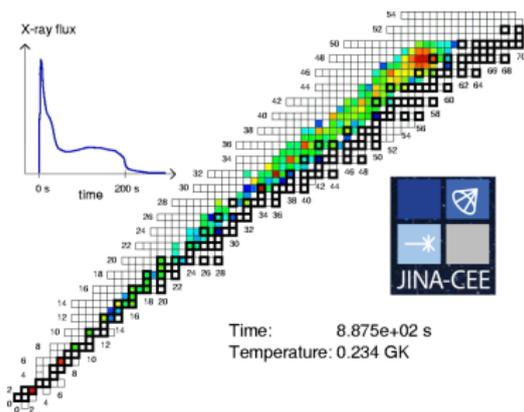


J.L. Tain et al., PRL 115 (2015) 062502
 E. Valencia et al., PRC 95 (2017) 024320

- 1 Introduction: TAGS technique
- 2 Astrophysical motivation
- 3 r-process
- 4 rp-process and X-ray bursts**
- 5 Summary

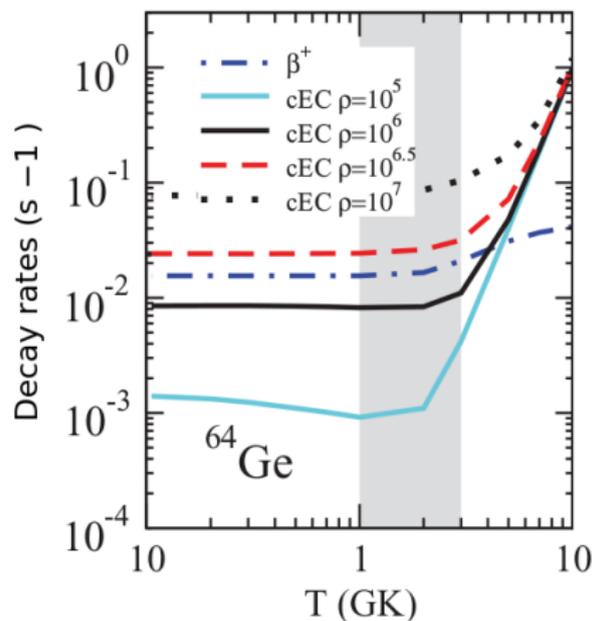
X-ray bursts: intense x-ray emission

- Type I: instabilities in the accretion disk (binary system: neutron star + main sequence or red giant star)
- $T_{peak} = 1-3$ GK and $\rho = 10^6-10^7$ g/cm³
- Light curves with a large variety of shapes
- Fast rise (1-10 s), peak luminosity of 10^{38} erg s⁻¹ + slower decline
- When β -decay competes with proton capture \Rightarrow waiting points



The role of electron capture

- Electron capture to weak-decay rates was traditionally neglected in model calculations of Type I X-ray bursts
- cEC process plays an important role in the weak-decay rates of nuclei close to the proton drip-line in XRB calculations.

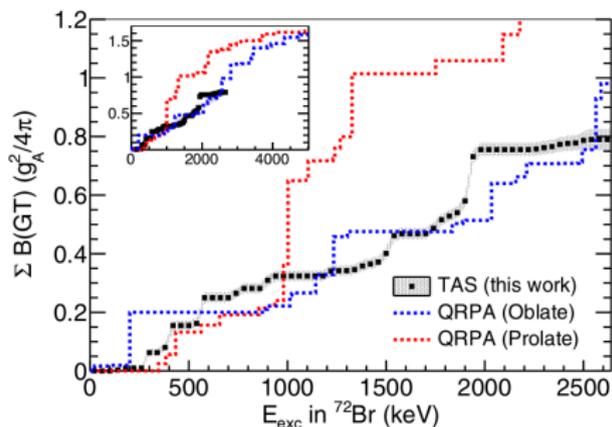


P. Sarriguren, PRC 83 (2011) 025801

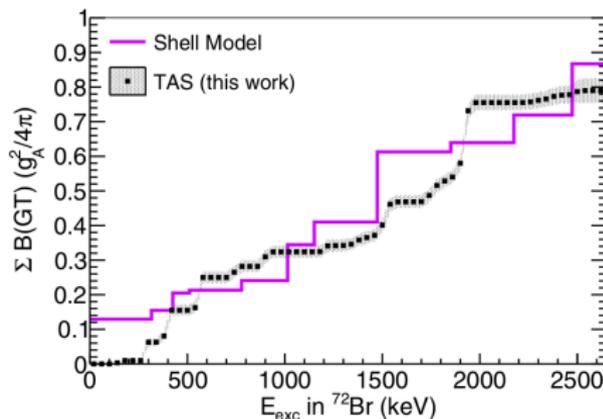
Waiting points and TAGS@ISOLDE: ^{72}Kr , ^{76}Sr

^{76}Sr : E. Nacher et al., PRL 92 (2004) 232501

^{72}Sr : J.A. Briz et al., PRC 92 (2015) 054326:



QRPA

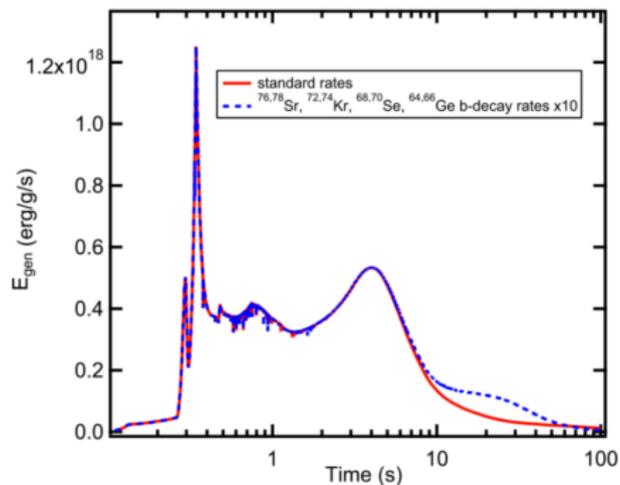


Shell Model

also beyond-mean-field: A. Petrovici a and O. Andrei EPJA (2015) 51: 133

Waiting points and TAGS@ISOLDE: ^{64}Ge and ^{68}Se

- IS570: $^{64,66}\text{Ge}$ in May 2015: analysis ongoing
- $^{68,70}\text{Se}$ foreseen after the long shutdown of CERN



← A. Parikh calculation
for the IS570 proposal

^{64}Ge and ^{68}Se : cEC-decay rates higher than the β^+ decay rates: factor 2
For their **N=Z+2 neighbours** factor 100

- 1 Introduction: TAGS technique
- 2 Astrophysical motivation
- 3 r-process
- 4 rp-process and X-ray bursts
- 5 Summary**

- TAGS results allow to compare experimental and theoretical β strength distributions

- TAGS results allow to compare experimental and theoretical β strength distributions
- Sensitivity to $I_{\beta\gamma}$ above S_n : larger than expected!

- TAGS results allow to compare experimental and theoretical β strength distributions
- Sensitivity to $I_{\beta\gamma}$ above S_n : larger than expected!
- Interaction of neutrons with the spectrometer

- TAGS results allow to compare experimental and theoretical β strength distributions
- Sensitivity to $I_{\beta\gamma}$ above S_n : larger than expected!
- Interaction of neutrons with the spectrometer
- Comparison with HF calculations: Γ_γ width $\rightarrow (n,\gamma)$

- TAGS results allow to compare experimental and theoretical β strength distributions
- Sensitivity to $I_{\beta\gamma}$ above S_n : larger than expected!
- Interaction of neutrons with the spectrometer
- Comparison with HF calculations: Γ_γ width $\rightarrow (n,\gamma)$
- Measurements at IGISOL and ISOLDE with nuclear astrophysics motivations

- TAGS results allow to compare experimental and theoretical β strength distributions
- Sensitivity to $I_{\beta\gamma}$ above S_n : larger than expected!
- Interaction of neutrons with the spectrometer
- Comparison with HF calculations: Γ_γ width $\rightarrow (n,\gamma)$
- Measurements at IGISOL and ISOLDE with nuclear astrophysics motivations
- Important decays for the r-process and for the rp-process

*Thank you very much for your
attention!*



UNIVERSITY OF
SURREY



JYVÄSKYLÄN YLIOPISTO
UNIVERSITY OF JYVÄSKYLÄ



BROOKHAVEN
NATIONAL LABORATORY